

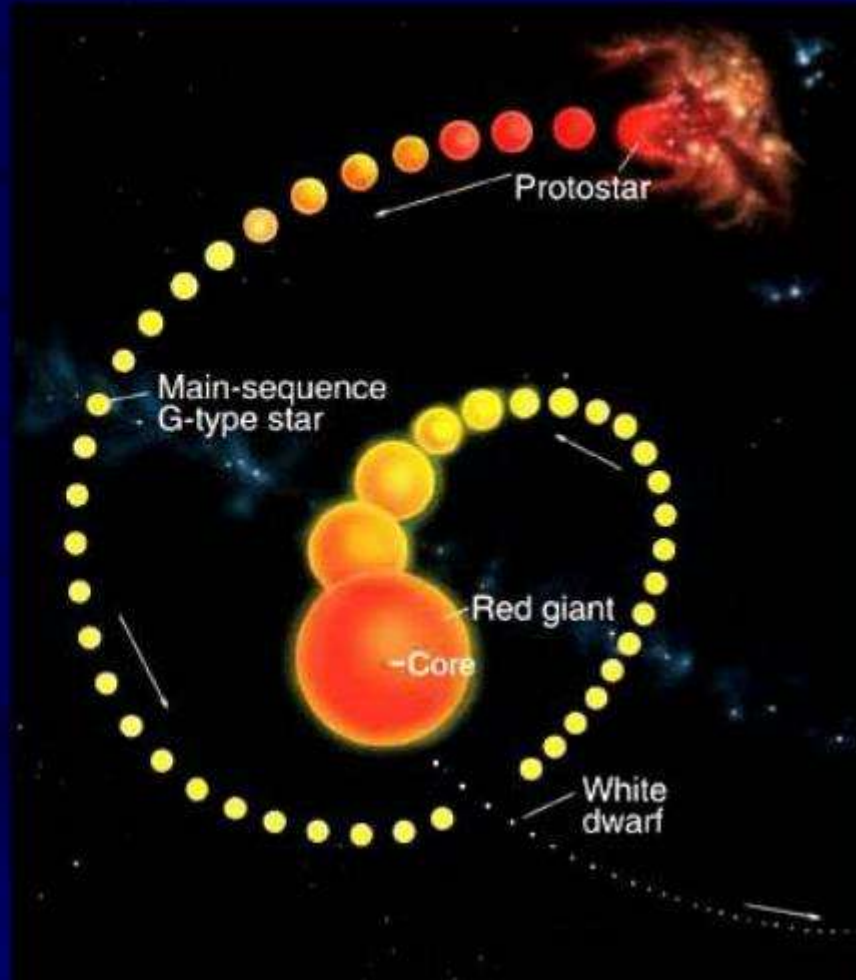
Fate of the Sun Fate of the Universe

Jim Rauf

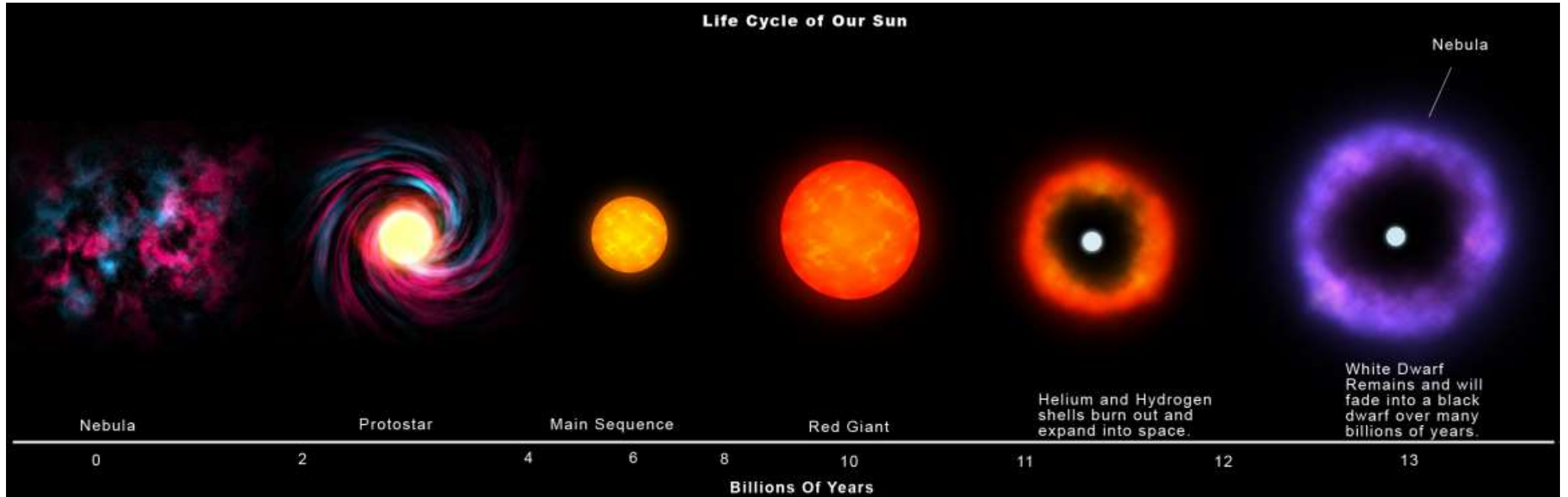
Lifecycle of The Sun -a Main Sequence G Star

Lifecycle

- Lifecycle of a main sequence G star
- Most time is spent on the main-sequence (normal star)

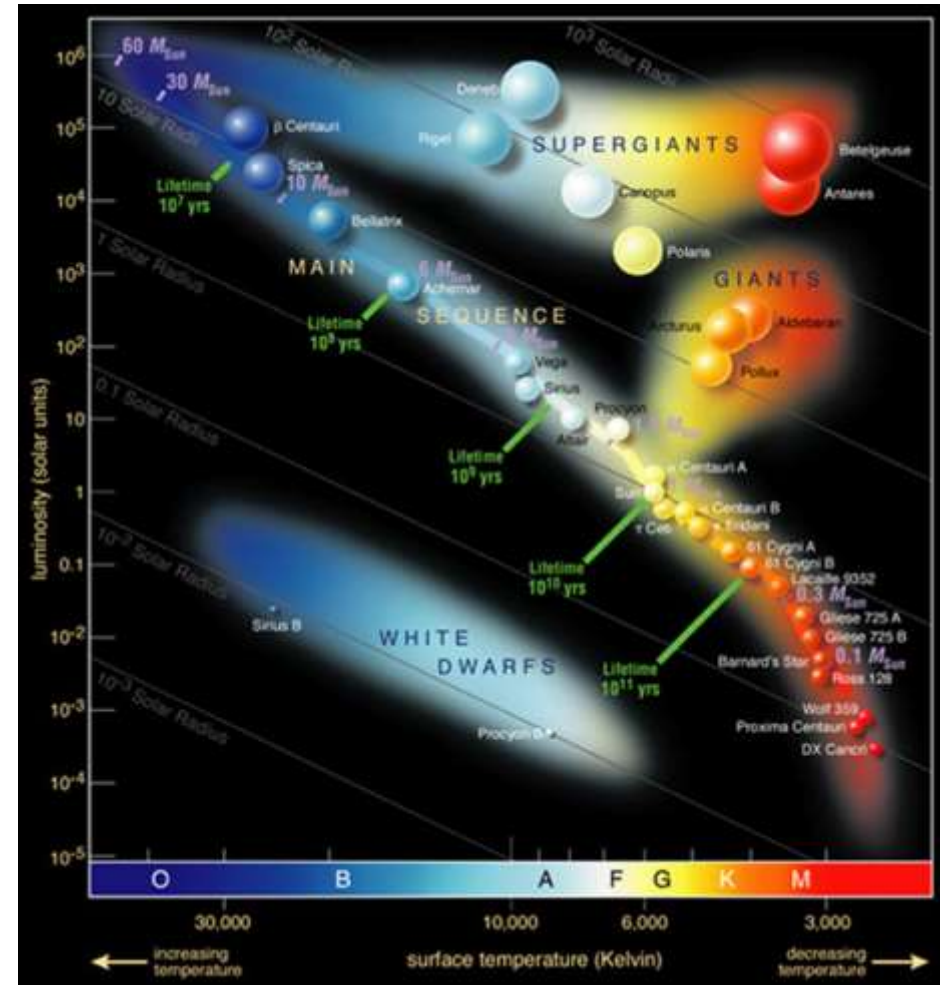


Fate of the Sun- Nebula to Nebula



Fate of the Sun

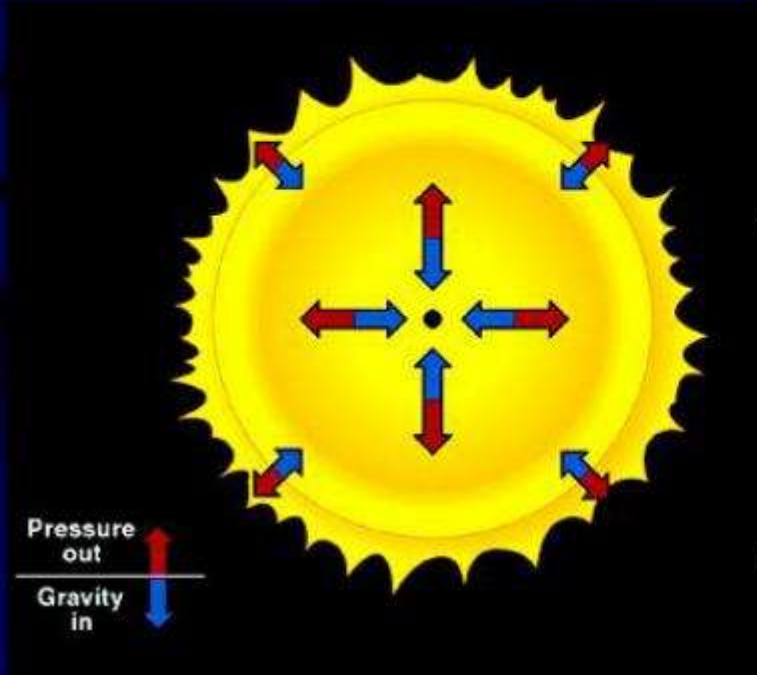
- The Main Sequence:
- The Sun, like most stars in the Universe, is on the **main sequence stage** of its life, during which nuclear fusion reactions in its core fuse hydrogen into helium
- Every second, 600 million tons of matter are converted into neutrinos, solar radiation, and roughly 4×10^{27} Watts of energy
- For the Sun, this process began 4.57 billion years ago, and it has been generating energy this way every since
- So far, the Sun has converted an estimated 100 times the mass of the Earth into **helium** and **solar energy**



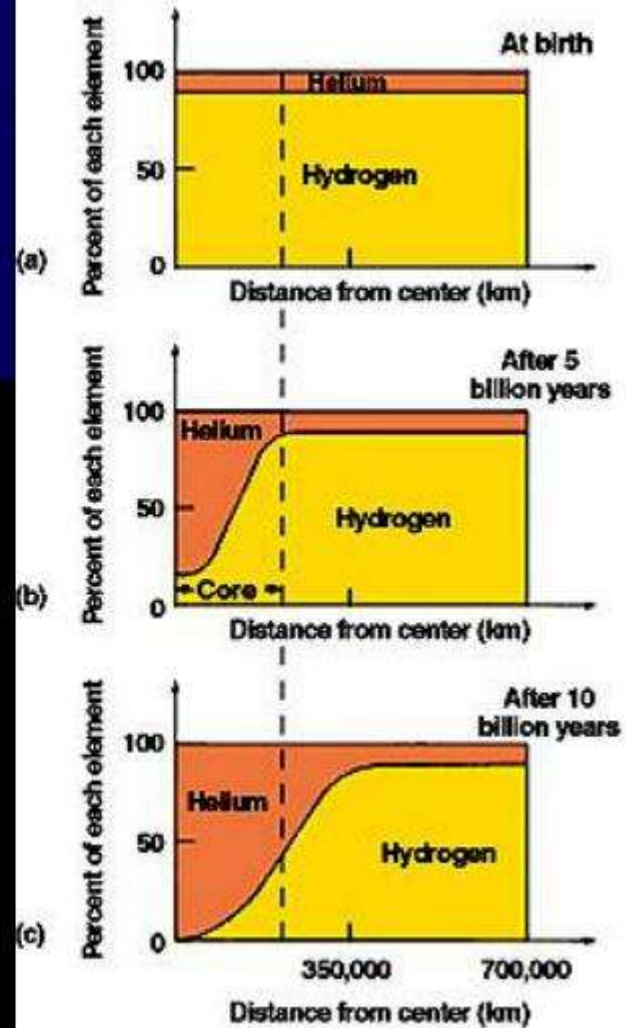
Fate of the Sun-Leaving the Main Sequence

Why Do Stars Leave the Main Sequence?

- Running out of fuel



The diagram illustrates the forces within a star. Red arrows point outwards from the center, representing the outward pressure. Blue arrows point inwards towards the center, representing the inward pull of gravity. A legend at the bottom left shows a red arrow pointing up labeled 'Pressure out' and a blue arrow pointing down labeled 'Gravity in'.



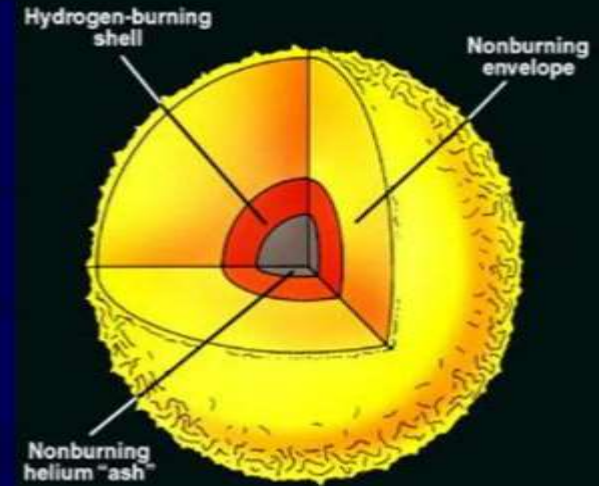
Fate of the Sun-Hydrogen Shell Burning

- As the supply of hydrogen in the core begins to decrease (having been fused into **helium**), the fusion rate goes down, and the amount of energy generated drops
- The temperature will then begin to drop and then the pressure will also decrease in the fusion core

$$PV = nRT$$

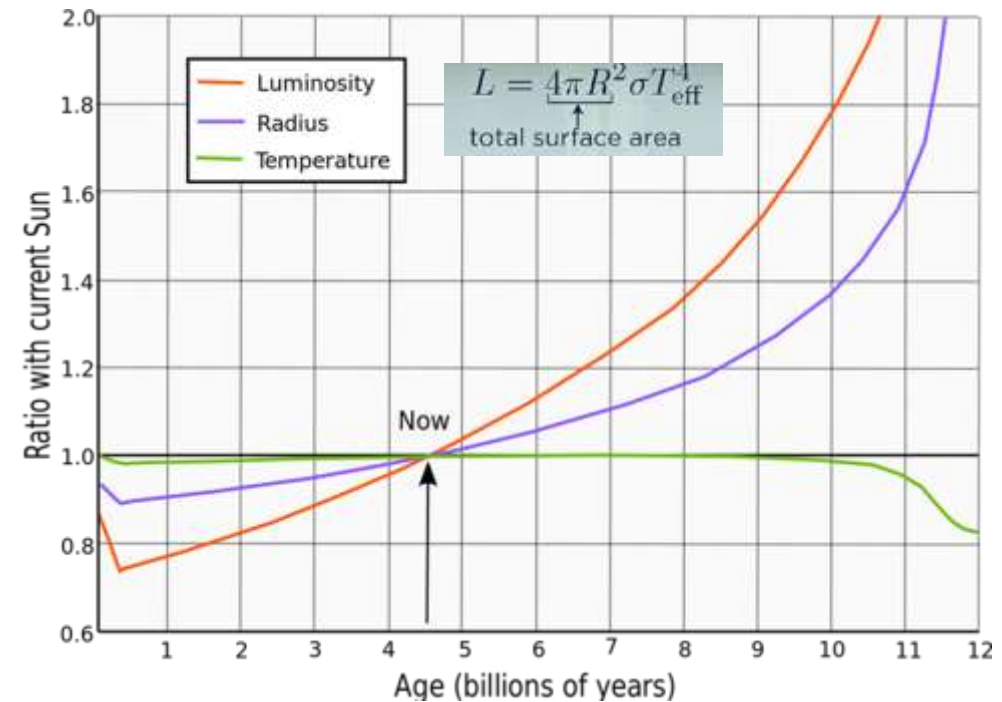
- A drop in pressure means that the core region of the star will contract slightly
- This will cause the temperature to go up again, and the fusion rate for the remaining hydrogen in the core increases
- The sharp rise in temperature also starts a **hydrogen burning shell** around the core, a region that before was too cool (less than 15 million K) to sustain fusion before

- Cooler core \rightarrow imbalance between pressure and gravity \rightarrow core shrinks
- hydrogen shell generates energy too fast \rightarrow outer layers heat up \rightarrow star expands
- Luminosity increases
- Duration \sim 100 million years
- Size \sim several Suns



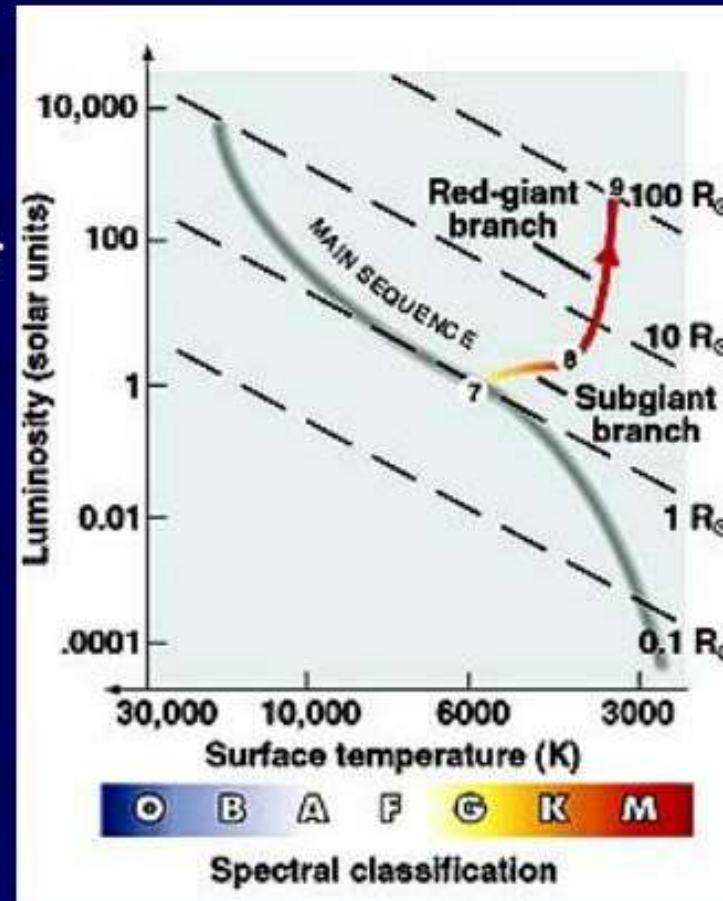
Fate of the Sun

- At present, this is leading to ~ 1% increase in luminosity every 100 million years
- And ~ 30% increase over the course of the last 4.5 billion years
- In 1.1 billion years from now, the Sun will be ~ 10% brighter than it is today, and this increase in luminosity will affect the planets in the **solar system**
- All the planets will receive more solar energy due to the luminosity increase
- In 3.5 billion years from now, the Sun will be 40% brighter than it is right now
- The Sun's increased luminosity (power output) will move the solar system's "habitable zone" outward



Fate of the Sun-Red Giant Phase

- Luminosity huge (~ 100 Suns)
- Surface Temperature lower
- Core Temperature higher
- Size ~ 70 Suns (orbit of Mercury)

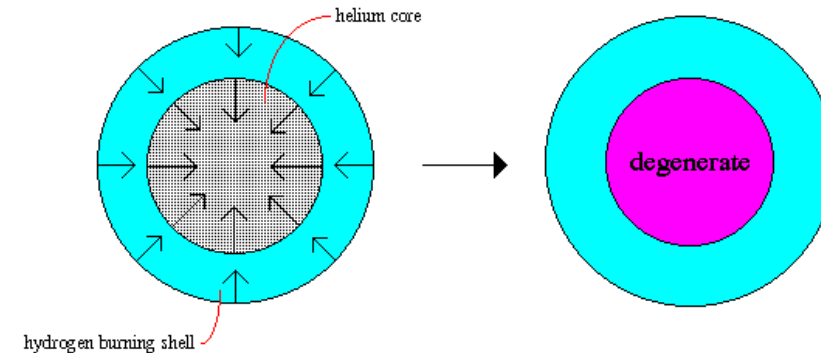


Fate of the Sun-After the Red Giant Phase

- A sun's evolution after the **red giant** phase depends on its mass
- The hydrogen burning shell eats its way outward leaving behind more **helium ash**
- The core becomes more massive and contracts
- Contraction heats the core, it becomes more dense
- The density of the core increases to where the electrons and helium nuclei become **degenerate**
- Electron and Neutron degeneracy are stellar applications of the **Pauli Exclusion Principle**
 - No two nucleons can occupy identical quantum states
- The core begins to act more like a liquid than a gas, it becomes incompressible
- Further contraction stops

Core Degeneracy

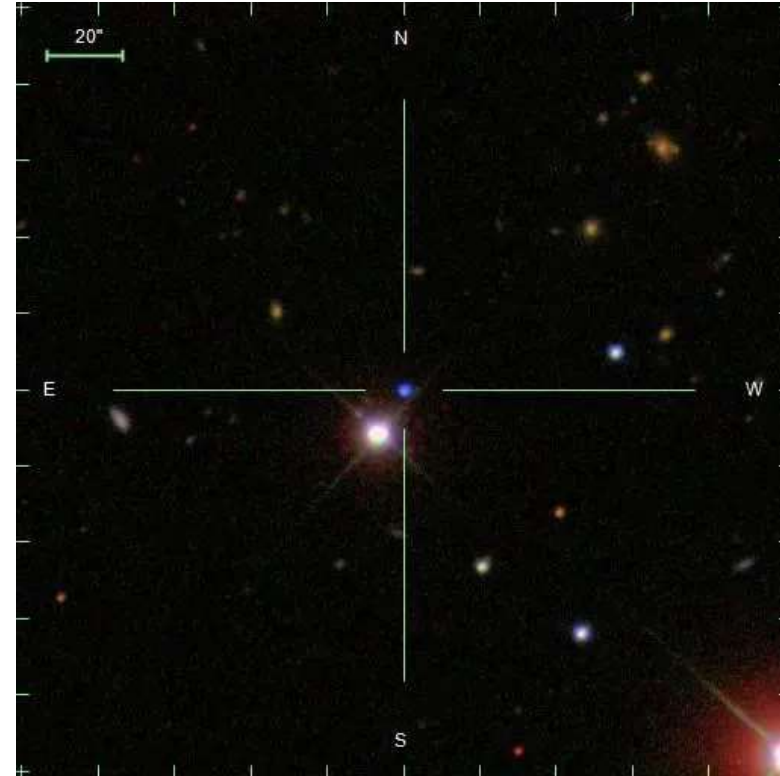
The hydrogen burning shell in red giants deposits helium ash into the core. The helium core increases in mass and contracts.



Contraction increases the pressure and density of the core until the electrons become degenerate. The core temperature goes up with no change in pressure until triple-alpha burning begins.

Fate of the Sun-After the Red Giant Phase

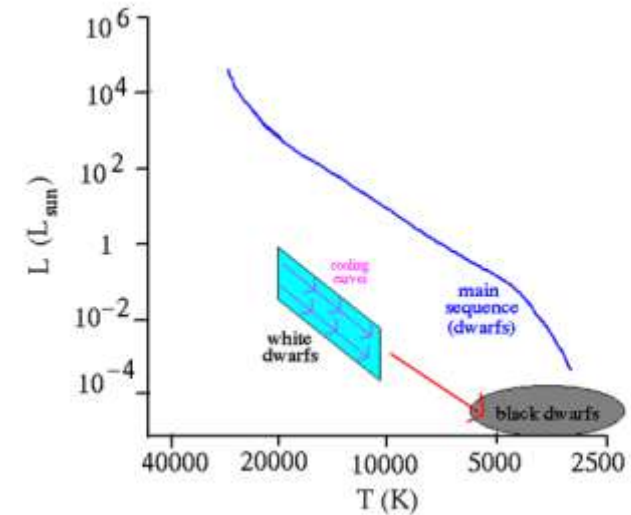
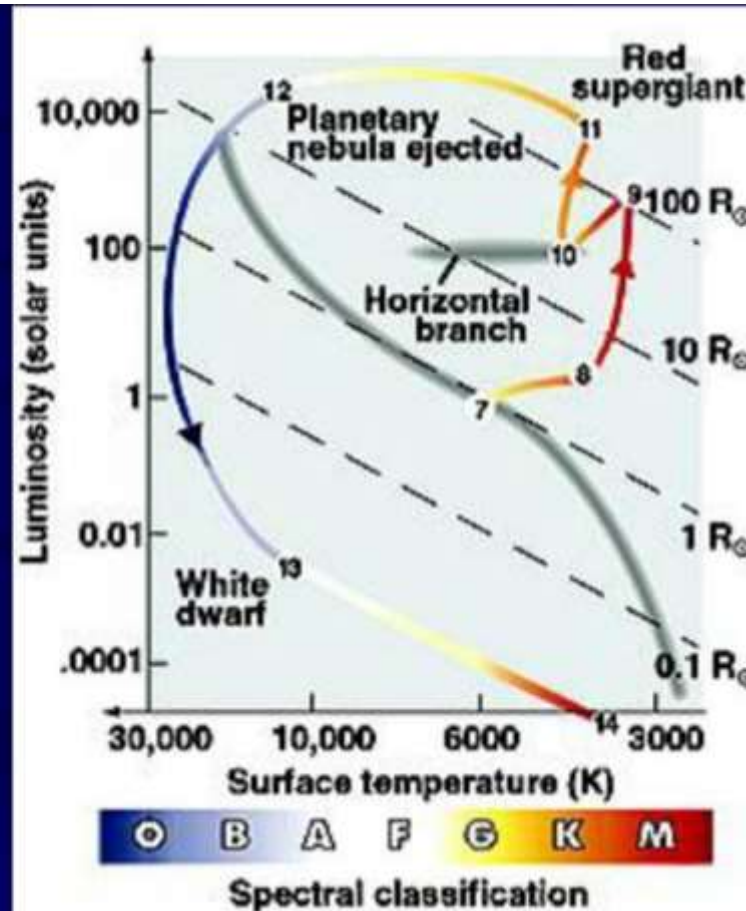
- The energy from the gravitational collapse is not sufficient to produce the neutrons (protons +electrons) of a neutron star
- Once the sun is degenerate, gravity cannot compress it any more
- Quantum mechanics dictates that there is no more available space to be taken upThe collapse is halted by electron degeneracy to form a **white dwarf**



White dwarf SDSS J124043.01+671034.68

Fate of the Sun -After the Red Giant Phase-White Dwarf

- Core radiates only by stored heat, not by nuclear reactions
- core continues to cool and contract
- Size ~ Earth
- Density: a million times that of Earth – 1 cubic cm has 1000 kg of mass!

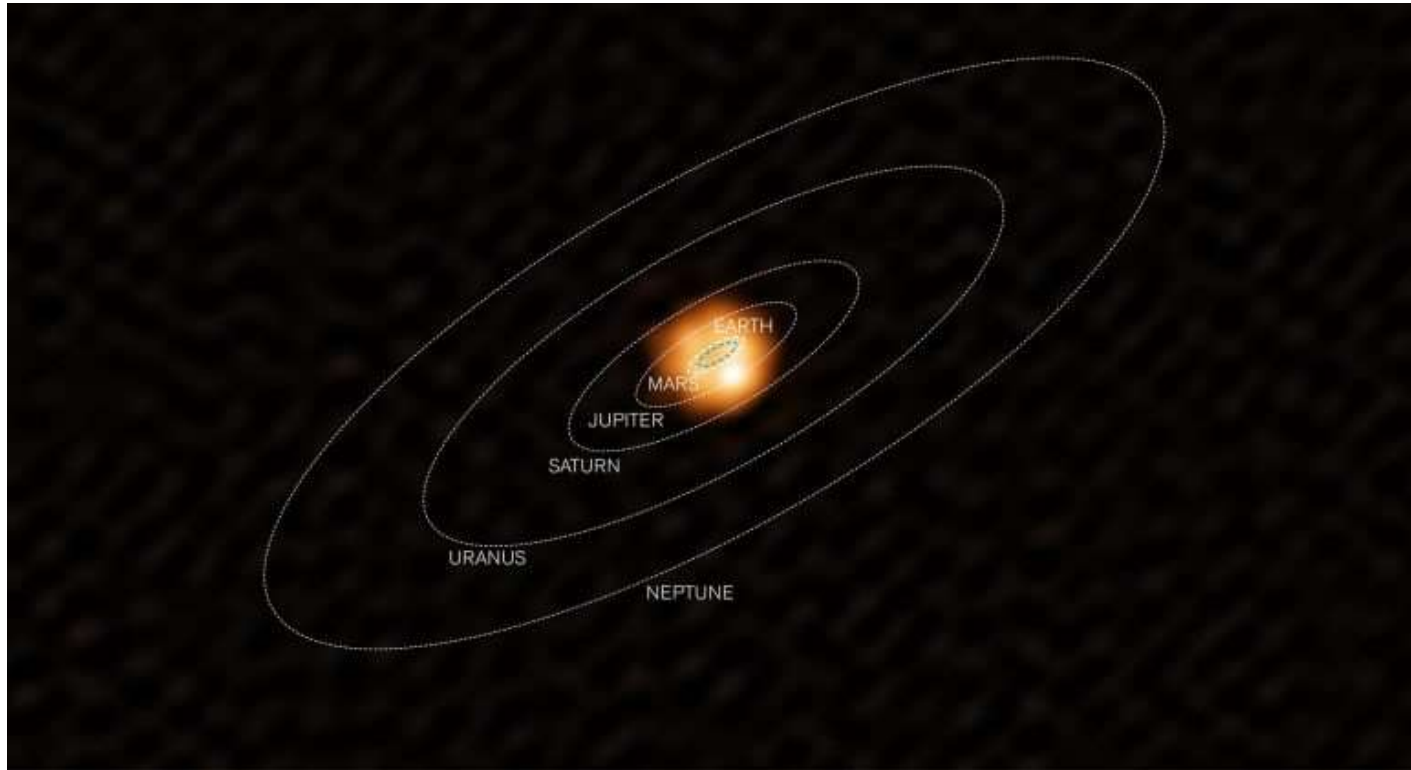


- The cooling process is very slow for white dwarfs
- After a billion years the white dwarf will be down to 0.001 the luminosity of the Sun
- The white dwarf will eventually give up all its energy and become a solid, crystal **black dwarf**

Fate of the Sun



Fate of the Sun



- New images show details on the surface of the red giant **W Hydrae**, 320 light years distant in the constellation of Hydra
- The star is a few billion years further on than the Sun in its life
- The dotted rings show the size of the orbits of the Earth (in blue) and other planets around the Sun for comparison

Fate of the Universe-Friedmann Equation –General Relativity

The Friedmann Equation

1. The “cosmological constant”

$$H^2 = \frac{2k}{a^2 r_0^2} + \frac{8\pi G}{3} \rho + \frac{\Lambda}{3}$$

2. Density includes both matter *and* energy

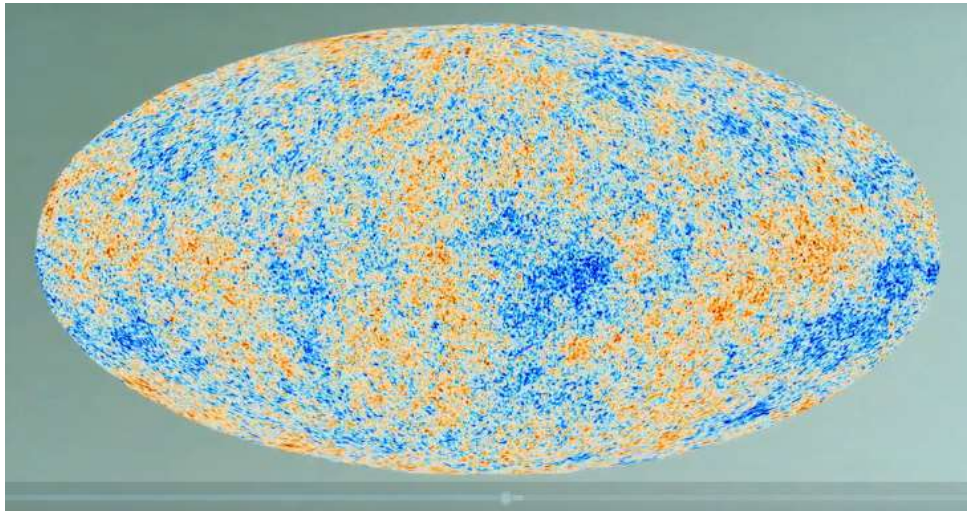
$$\rho \longrightarrow \rho + \frac{u}{c^2}$$

3. Energy/mass k \longrightarrow curvature of space

- $K=0$ Space is Flat
- $K \neq 0$ Space is Curved

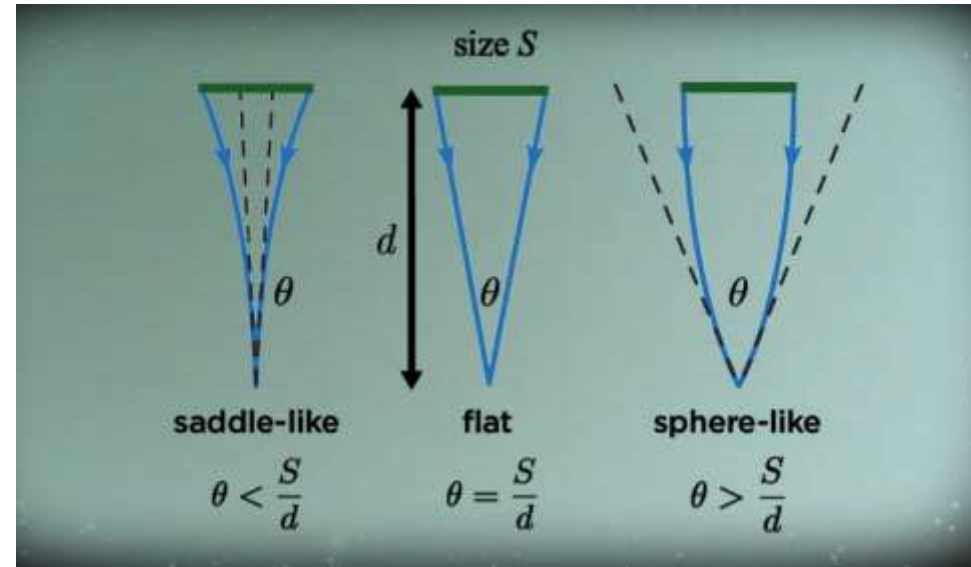
- $K < 0$ Space is Curved Like a Sphere
- $K \geq 0$ Space is Curved Like a Saddle

Fate of the Universe-Friedmann Equation –General Relativity




Cosmic Microwave Background Radiation

Theta = S/d --- $K=0$ and Universe is flat



Measure dimensions of “hot spots” in CMB

Fate of the Universe-Friedmann Equation –General Relativity

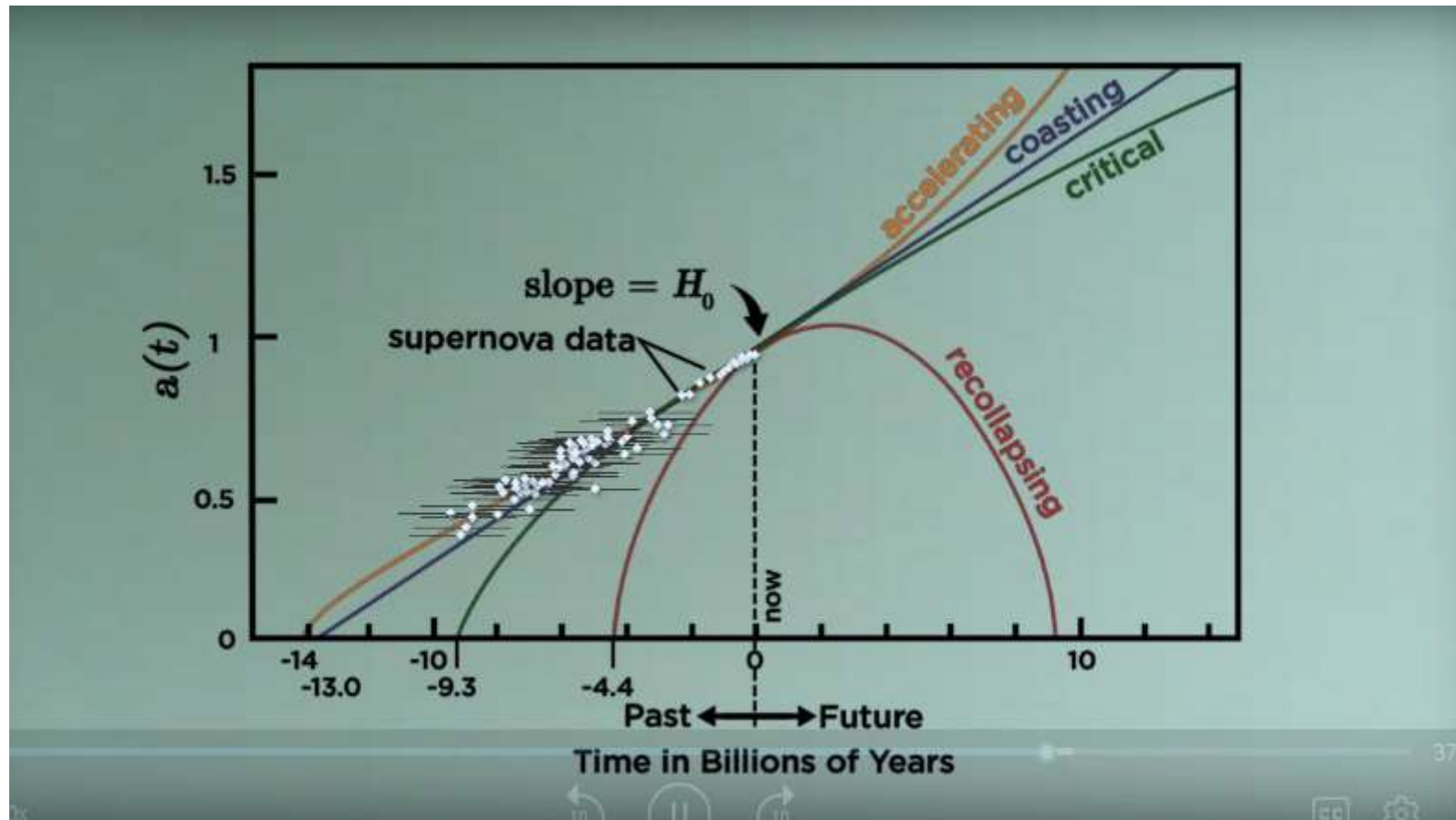


1. The universe will expand forever
2. The age of the universe < 9.3 billion years
3. Space is curved like a Pringle

Problems

1. Some star clusters are 13 billion years old
2. CMB suggests space is flat, not curved

Fate of the Universe-Friedmann Equation –General Relativity



The universe is expanding at an accelerating rate

Fate of the Universe-Friedmann Equation –General Relativity

$$H^2 = \frac{8\pi G}{3} \left(\rho + \frac{\Lambda}{8\pi G} \right)$$

↑ 30% of critical density
↑ 70% of critical density

Density Dark Energy

The universe will continue to expand at an accelerating rate

